# One fish, two fish, red fish, blue fish



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# The Riemann problem at infinite d.



A. Karch, H. C. Chang, I. Amado, C. Herzog. M. Spillane



Cold

Hot

inhomogenous



Cold

Hot



Cold

Hot



The evolution is governed by the dynamics of a conformal field theory:

$$\partial_{\mu}T^{\mu\nu} = 0 \qquad T^{\mu}{}_{\mu} = 0$$





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Bernard & Doyon, 2012
Bhaseen et. al. 2013
H. -C. Chang, Karch & AY et. al. 2013
Amado & AY, 2015
Megias, 2015
PourHasan, 2015
Bachas, Skenderis & Withers, 2015
Lucas et. al., 2015
Herzog & Spillane, 2015
Herzog, Spillane, AY, 2016

Plan:

- •Two dimensional systems
- •Discussion of d>2
- •Large d.













In a conformal theory

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#### Equivalently

$$\partial_t T^{tt} = -\partial_x T^{tx} \qquad \partial_t T^{tx} = -\partial_x T^{xx} \qquad -T^{tt} + T^{xx} = 0$$



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whose solution is

$$T^{\mu\nu} = \begin{pmatrix} T_+(t+x) + T_-(-t+x) & T_-(-t+x) - T_+(t+x) \\ T_-(-t+x) - T_+(t+x) & T_+(t+x) + T_-(-t+x) \end{pmatrix}$$



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The initial conditions imply that:

$$T_{-}(u) = T_{+}(u), \quad 2T_{+}(u) = T^{11}(t = 0, x = u), \quad 0 < u < L$$



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The initial and boundary conditions imply that:

$$T_{-}(u) = T_{+}(u), \quad 2T_{+}(u) = T^{11}(t = 0, x = u), \qquad 0 < u < L$$
  
$$T_{-}(-u) + T_{+}(u) = P_{\text{left}}, \quad T_{-}(-u + L) + T_{+}(u + L) = P_{\text{right}}, \qquad \forall u$$



# We find: $T_{+}(u) = \begin{cases} -n(P_{\text{left}} - P_{\text{right}}) + T_{+}(u_{0}) & u_{0} > 0\\ -(n-1)(P_{\text{left}} - P_{\text{right}}) + (P_{\text{right}} - T_{-}(-u_{0})) & u_{0} < 0 \end{cases}$ $T_{-}(u) = \begin{cases} -n(P_{\text{left}} - P_{\text{right}}) + T_{-}(u_{0}) & u_{0} > 0\\ -(n-1)(P_{\text{left}} - P_{\text{right}}) + (Pn_{\text{right}} - T_{+}(-u_{0})) & u_{0} < 0 \end{cases}$

where

 $u = u_0 + 2nL \qquad -L < u_0 < L \qquad n \in \mathbb{Z}$ 



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If we set  $L \rightarrow \infty$ , then in the finite x,  $t \rightarrow \infty$  limit we expect to see a time independent steady state.

### Recall that in a conformal theory:

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The exact same analysis can be used to consider more complicated configurations:



**F**01

X

The exact same analysis can be used to consider more complicated configurations:

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## Main ingredient:

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$$T^{\mu\nu}(t,x) = \begin{pmatrix} T^{00} & T^{01} & 0\\ T^{01} & T^{11} & 0\\ 0 & 0 & T_{\perp} \end{pmatrix}$$

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We need more input.

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Let us assume, in addition, that the system is described by a perfect inviscid fluid:

$$T^{\mu\nu} = \epsilon(P)u^{\mu}u^{\nu} + (\eta^{\mu\nu} + u^{\mu}u^{\nu})P$$

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 $T^{\mu\nu} = \epsilon(P) u^{\mu} u^{\nu} + (\eta^{\mu\nu} + u^{\mu} u^{\nu}) P$ For a conformal field theory:

$$\epsilon = (d-1)P$$

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The resulting steady state will be characterised by:  $T_*^{00}$   $T_*^{01}$ 







#### Let us move to a frame where s=0, we obtain, e.g.,

#### "Good" shocks





Energy momentum conservation and conformal invariance imply:

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The resulting system of non linear equations is still difficult to solve.

Consider another ansatz (rarefaction wave)

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Consider another ansatz (rarefaction wave)

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Up to an overall rescaling, we have a three parameter family of solutions.





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- Is this correct?





Let us start by considering an equilibrated configuration



A planar event horizon:



$$ds^2 = 2dt \left(dr - A(r)dt\right) + r^2 d\vec{x}^2$$



$$as' = 2at \left(ar - A(r)at\right) + r' ax$$
$$A(r) = r^2 \left(1 - \left(\frac{4\pi T}{3r}\right)^3\right)$$



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A planar event horizon:  $ds^2 = 2dt \left(dr - A(r)dt\right) + r^2 d\vec{x}^2$  $P(T) = p_0 \left(\frac{4\pi T}{3}\right)^3$  $A(r) = r^2 \left( 1 - \left(\frac{4\pi T}{3r}\right)^3 \right)$ e.g., in ABJM  $p_0 = \frac{2N^2}{0\sqrt{2\lambda}} \qquad \lambda = \frac{N}{k}$ 



$$P(T_L) = p_0 \left(\frac{4\pi T_L}{3}\right)^3$$

$$P(T_R) = p_0 \left(\frac{4\pi T_R}{3}\right)^3$$





$$A(r,z) = r^2 \left( 1 - \left(\frac{a_1(z)}{3r}\right)^3 \right)$$
Out of equilibrium we want to start with:



$$a_1(-\infty) = \frac{4\pi T_L}{3}$$
$$a_1(\infty) = \frac{4\pi T_R}{3}$$

Out of equilibrium we want to start with:

$$ds^{2} = 2dt \left(dr - A(r, z)dt\right) + r^{2}d\vec{x}^{2}$$

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#### and evolve it forward in time. Using

$$ds^{2} = 2dt(dr - A(t, z, r)dt - F(t, z, r)dz) + \Sigma^{2}(t, r, z)\left(e^{B(t, z, r)}dx_{\perp}^{2} + e^{-B(t, z, r)}dz^{2}\right)$$

the Einstein equations reduce to a set of nested linear differential equations in the radial coordinate 'r'.

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the Einstein equations reduce to a set of nested linear differential equations in the radial coordinate 'r'. We have solved these equations numerically.

(Chesler, Yaffe, 2012)













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with the scaling:

$$\epsilon = \mathcal{O}(d^0) \qquad \qquad x = \frac{\chi}{\sqrt{d}} = \mathcal{O}(1/\sqrt{d})$$
$$x_{\perp} = \frac{\chi_{\perp}}{d} = \mathcal{O}(1/\sqrt{d}) \qquad v = \frac{\beta}{\sqrt{d}} = \mathcal{O}(1/\sqrt{d})$$

If we send d to infinity the difference between the solutions becomes prominent.

For the ideal inviscid fluid,

$$T^{\mu\nu} = \frac{\epsilon}{d-1} \left( du^{\mu}u^{\nu} + \eta^{\mu\nu} \right)$$

with the scaling:

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one finds (  $e = \epsilon$ ,  $j = \epsilon\beta$ )
$$T^{\mu\nu} = \begin{pmatrix} e & j \\ j & e + \frac{j^2}{e} \end{pmatrix}$$

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one finds (  $e = \epsilon$ ,  $j = \epsilon\beta$  )

$$T^{\mu\nu} = \begin{pmatrix} e & j \\ j & e + \frac{j^2}{e} \end{pmatrix}$$

Consider

$$e(\chi, t) = (e_r - e_l)\theta(x - st) + e_l$$
$$j(\chi, t) = (j_r - j_l)\theta(x - st) + j_l$$

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If we choose a reference frame with  $j_l = 0$ and scale the energy so that  $e_l = 1$ then we have a one parameter family of solutions  $e_r(s) \quad j_r(s)$ 

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-0.4









If we send d to infinity the difference between the solutions becomes prominent.

One finds

$$T^{\mu\nu} = \begin{pmatrix} e & j \\ j & e + \frac{j^2}{e} \end{pmatrix}$$

#### Consider

$$e(\chi, t) = e(\chi/t)$$
  
 $j(\chi, t) = j(\chi/t)$   
with  $j_l = 0$  and  $e_l = 1$ 













We can fix instead the right asymptotic:  $e_r = 1$   $j_r = 0$ 



Or:



Or:


Or:

















We can now glue solutions like we did before:

RS  $e_{l,e_r}$  $o_{d}$  $o_{d}$ 

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#### We obtain a "phase" diagram,



Out of equilibrium we would like to solve the equations of motion for:

 $ds^{2} = dt(2dr - g_{tt}dt - 2g_{t\chi}d\chi) + g_{\chi\chi}d\chi^{2} + g_{\perp\perp}d\chi_{\perp}^{2}$ 

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(Emparan, Suzuki, Tanabe, 2015 (see also Bhattacharyya et. al. 2015))

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$$\frac{g_{\chi\chi}}{r^2} = \frac{1}{d} + \mathcal{O}(d^{-2})$$

$$\frac{g_{\perp\perp}}{r^2} = \frac{1}{d} + \mathcal{O}(d^{-3})$$

where:

$$\partial_t e - \partial_\chi^2 e = -\partial_\chi j \qquad \qquad \partial_t j - \partial_\chi^2 j = -\partial_\chi \left(\frac{j^2}{e} + e\right)$$

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which come from a conservation of

$$T^{\mu\nu} = \begin{pmatrix} e & j - \partial_{\chi} e \\ j - \partial_{\chi} e & e + \frac{j^2}{e} - 2\partial_{\chi} j + \partial_{\chi}^2 e \end{pmatrix} + \begin{pmatrix} \partial_{\chi}^2 g & -\partial_{\chi} \partial_t g \\ -\partial_{\chi} \partial_t g & \partial_t^2 g \end{pmatrix}$$

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## Summary



### Summary



#### In a 2d CFT we find

$$T^{00} = T_{+}(\infty) + T_{-}(-\infty) = \frac{1}{2} \left( P_{\text{left}} + P_{\text{right}} \right) ,$$
$$T^{01} = T_{-}(-\infty) - T_{+}(\infty) = \frac{1}{2} \left( P_{\text{left}} - P_{\text{right}} \right)$$

d>2

#### We have 4 possibilities:







# Thank you